## AN EXTENSION OF THE VEKUA-BITSADZE METHOD FOR SOLVING EQUATIONS OF THE SHELLS

## Meunargia T.

I. Vekua Institute of Applied Mathematics of Iv. Javakhishvili Tbilisi State University

*Key words*: Shell, Vekua-Bitsadze method parameter, nonlinear theory. *MSC 2000*: 74K25.

I. Vekua has constructed several versions of the refined linear theory of thin and shallow shells, containing the regular process by means of the method of reduction of three-dimensional problems of elasticity to two-dimensional ones.

In the present paper by means of the I. Vekua method the system of differential equations for the nonlinear theory of non-shallow shells is obtained. Using the method I. Vekua and the method of a small parameter 2-D system of equations for the nonlinear and non-shallow shells is obtained. For any approximations of order N the complex representations of Vekua-Bitsadze type [2] of the general solutions are obtained.

Under thin and shallow shells I. Vekua meant three-dimensional shell-type elastic bodies satisfying the conditions

$$a_{\alpha}^{\beta} - x_3 b_{\alpha}^{\beta} \cong a_{\alpha}^{\beta}, \ -h(x^1, x^2) \le x_3 \le h(x^1, x^2) \ (\alpha, \beta = 1, 2),$$
 (\*)

where  $a_{\alpha}^{\beta}$  and  $b_{\alpha}^{\beta}$  are mixed components of the metric tensor and the curvature tensor of the shell's midsurface,  $x_3$  is the thickness coordinate and h is the semi-thickness, depending on curvilinear coordinates  $x^1, x^2$ .

In the sequel, under by non-shallow shells we mean elastic bodies not subject to assumption (\*), i. e., such that

$$a_{\alpha}^{\beta} - x_3 b_{\alpha}^{\beta} \not\cong a_{\alpha}^{\beta} \Rightarrow |x_3 b_{\alpha}^{\beta}| \le q < 1.$$

1. To construct the theory of shells we use the coordinate system which is normally connected with the midsurface S. This means that the radius-vector of any point of the domain  $\Omega$  can be represented in the form [1]

$$\mathbf{R}(x^1, x^2, x^3) = \mathbf{r}(x^1, x^2) + x^3 \mathbf{n}(x^1, x^2) \ (x^3 = x_3),$$

where  $\boldsymbol{r}$  and  $\boldsymbol{n}$  are radius-vector and the unit vector of the normal of the midsurface  $S(x_3 = 0)$ ;  $x^1, x^2$  are the Gaussian parameters of S.

Covariant and contravariant basis vectors  $\mathbf{R}_i$  and  $\mathbf{R}^i$  of the surface  $\hat{S}(x_3 = \text{const})$  and the corresponding basis vectors  $\mathbf{r}_i$  and  $\mathbf{r}^i$  of the midsurface  $S(x_3 = 0)$  are connected by the following relations [1]:

$$\mathbf{R}_{i} = A_{i}^{\ j} \mathbf{r}_{j} = A_{ij} \mathbf{r}^{j}, \ \mathbf{R}^{i} = A_{\ j}^{i} \mathbf{r}^{j} = A^{ij} \mathbf{r}_{j} \ (i, j = 1, 2, 3),$$

where

$$A_{\alpha.}^{\beta} = a_{\alpha}^{\beta} - x_{3}b_{\alpha}^{\beta}, \quad A_{\beta}^{\alpha} = \vartheta^{-1}[(1 - 2Hx_{3})a_{\beta}^{\alpha} + x_{3}b_{\beta}^{\alpha}], \quad A_{3}^{i} = A_{i}^{3} = \delta_{i}^{3}, \\ \vartheta = 1 - 2Hx_{3} + Kx_{3}^{2}, \quad \mathbf{R}_{3} = \mathbf{R}^{3} = \mathbf{r}_{3} = \mathbf{r}^{3} = \mathbf{n} \quad (\alpha, \beta = 1, 2).$$
(1)

Here  $(a_{\alpha\beta}, a^{\alpha\beta}, a^{\alpha}_{\beta})$  and  $(b_{\alpha\beta}, b^{\alpha\beta}, b^{\alpha}_{\beta})$  are the components (co, contra, mixed) of the metric tensor and curvature tensor of the midsurface S. By H and K we denote a middle and Gaussian curvature of the surface S, where

$$2H = b_{\alpha}^{\alpha} = b_1^1 + b_2^2, \quad K = b_1^1 b_2^2 - b_2^1 b_1^2.$$

The main quadratic forms of the midsurface S have the form

$$I = ds^2 = a_{\alpha\beta} dx^{\alpha} dx^{\beta}, \quad II = k_s ds^2 = b_{\alpha\beta} dx^{\alpha} dx^{\beta}, \quad (2)$$

where  $k_s$  is the normal curvature of the surface S, and

$$a_{\alpha\beta} = \boldsymbol{r}_{\alpha}\boldsymbol{r}_{\beta}, \ b_{\alpha\beta} = -\boldsymbol{r}_{\alpha}\boldsymbol{n}_{\beta}, \ k_s = b_{\alpha\beta}s^{\alpha}s^{\beta}, s^{\alpha} = \frac{dx^{\alpha}}{ds}.$$

Here and in the sequel, under a repeated indices we mean summation; note that the Greek indices range over 1, 2, while Latin indices range over 1, 2, 3.

To construct the theory of non- shallow shells, it is necessary to obtain formulas for a family of surfaces  $\hat{S}(x_3 = \text{const})$ , analogous to (2) of the midsurface  $S(x_3 = 0)$  which have the form [1]

$$\mathbf{I} = d\hat{s}^2 = g_{\alpha\beta} dx^{\alpha} dx^{\beta}, \quad \mathbf{II} = k_{\hat{s}} d\hat{s}^2 = \hat{b}_{\alpha\beta} dx^{\alpha} dx^{\beta}, \quad (3)$$

where

$$g_{\alpha\beta} = a_{\alpha\beta} - 2x_3b_{\alpha\beta} + x_3^2(2Hb_{\alpha\beta} - Ka_{\alpha\beta}), \ \hat{b}_{\alpha\beta} = (1 - 2Hx_3)b_{\alpha\beta} + x_3Ka_{\alpha\beta},$$

and  $k_{\hat{s}}$  the normal curvature of the surface  $\hat{S}$ .

It is not now difficult to get the expression for the tangential normal  $\hat{l}$  of the surface  $\hat{S}$  directed to  $\hat{s}$  [3]:

$$\hat{\boldsymbol{l}} = \hat{\boldsymbol{s}} \times \boldsymbol{n} = [(1 - x_3 k_s) \boldsymbol{l} - x_3 \tau_s \boldsymbol{s}] \frac{ds}{d\hat{s}}, \quad d\hat{s} = \sqrt{1 - 2x_3 k_s + x_3^2 (k_s^2 + \tau_s^2)} ds,$$

where s and l are the unit vectors of the tangent and tangential normal on S,  $d\hat{s}$  and ds are the linear elements of the surfaces  $\hat{S}$  and S, and  $\tau_s$  is the geodesic torsion of the surface S.

2. We write the equation of equilibrium of elastic shell-type bodies in a vector form

$$\frac{1}{\sqrt{g}}\frac{\partial\sqrt{g}\boldsymbol{\sigma}^{i}}{\partial x^{i}} + \Phi = 0, \Rightarrow \nabla_{i}\boldsymbol{\sigma}^{i} + \Phi = 0, \qquad (4)$$

where g is the discriminant of the metric quadratic form of the three-dimensional domain  $\Omega$ ,  $\nabla_i$  are covariant derivatives with respect to the space coordinates  $x^i$ ,  $\Phi$  is on external force,  $\sigma^i$  are the contravariant constituents of the stress vector  $\boldsymbol{\sigma}_{(l)}^*$  acting on the area with the normal  $\boldsymbol{l}$  and representable as the Cauchy formula as follows:

$$\boldsymbol{\sigma}_{(l)} = \boldsymbol{\sigma}^{i} \overset{*}{l}_{i}, \quad \left( \overset{*}{l}_{i} = \overset{*}{l} \boldsymbol{R}_{i}, \right).$$

For the stress vector acting on the area with normal  $\hat{l}$ , we obtain

$$\boldsymbol{\sigma}_{(\hat{\boldsymbol{l}})} = \boldsymbol{\sigma}^{\alpha}(\hat{\boldsymbol{l}}\boldsymbol{R}_{\alpha}) = \vartheta \boldsymbol{\sigma}^{\alpha}(\boldsymbol{l}\boldsymbol{r}_{\alpha})\frac{ds}{d\hat{s}}.$$
 (5)

The stress-strain relation for the geometrically nonlinear theory of elasticity has the form

$$\boldsymbol{\sigma}^{i} = \sigma^{ij}(\boldsymbol{R}_{j} + \partial_{j}\boldsymbol{U}) = E^{ijpq}e_{pq}(\boldsymbol{R}_{j} + \partial_{j}\boldsymbol{U}), \qquad (6)$$

where  $\sigma^{ij}$  are contravariant components of the stress tensor,  $e_{ij}$  are covariant components of the strain tensor, **U** is the displacement vector,  $E^{ijpq}$  and  $e_{ij}$  are defined by the formulas:

$$E^{ijpq} = \lambda g^{ij} g^{pq} + \mu (g^{ip} g^{jq} + g^{iq} g^{ip}), \ e_{ij} = \frac{1}{2} (\boldsymbol{R}_i \partial_j \boldsymbol{U} + \boldsymbol{R}_j \partial_i \boldsymbol{U} + \partial_i \boldsymbol{U} \partial_j \boldsymbol{U}).$$
(7)

To reduce the three-dimensional problems of the theory of elasticity to the two-dimensional problems, it is necessary to rewrite the relation (4-7) in forms of the bases of the midsurface S of the shell  $\Omega$ .

The relation (4) can be written as

$$\frac{1}{\sqrt{a}}\frac{\partial\sqrt{a}\vartheta\boldsymbol{\sigma}^{\alpha}}{\partial x^{\alpha}} + \frac{\partial\vartheta\boldsymbol{\sigma}^{3}}{\partial x^{3}} + \vartheta\Phi = 0, \quad (a = a_{11}a_{22} - a_{12}^{2}). \tag{8}$$

From (1), (6), (7) we obtain

$$\boldsymbol{\sigma}^{i} = A_{i_{1}}^{i} A_{p_{1},}^{p} M^{i_{1}j_{1}p_{1}q_{1}} [(\boldsymbol{r}_{q_{1}}\partial_{p}\boldsymbol{U}) + \frac{1}{2} A_{q_{1}}^{q} (\partial_{p}\boldsymbol{U}\partial_{q}\boldsymbol{U})](\boldsymbol{r}_{j_{1}} + A_{j_{1}}^{j}\partial_{j}\boldsymbol{U}), \quad (9)$$
$$M^{i_{1}j_{1}p_{1}q_{1}} = \lambda a^{i_{1}j_{1}} a^{p_{1}q_{1}} + \mu (a^{i_{1}p_{1}}a^{j_{1}q_{1}} + a^{i_{1}q_{1}}a^{j_{1}p_{1}}) \quad (a^{ij} = \boldsymbol{r}^{i}\boldsymbol{r}^{j})$$

**3.** The isometric system of coordinates on the surface S is of special interest, for in this system we can obtain basic equations of the theory of shells in a complex form which in turn allows one to construct for a rather wide class of problems complex representations of general solutions by means of analytic functions of one variable  $z = x^1 + ix^2$ .

The main quadratic forms in the system of coordinates are of the type

$$\mathbf{I} = ds^2 = \Lambda(z, \bar{z}) dz d\bar{z}, \quad \mathbf{II} = k_s ds^2 = \frac{1}{2} \Lambda \left[ \bar{Q} dz^2 + 2H dz d\bar{z} + Q d\bar{z}^2 \right],$$

$$Q = 0.5(b_1^1 - b_2^2 + 2ib_2^1), \quad \Lambda(z, \bar{z}) > 0.$$

Introducing the well-known differential operators  $2\partial z = \partial_1 - i\partial_2$ ,  $2\partial_{\bar{z}} = \partial_1 + i\partial_2$  and the notations  $\tau_{.j}^{i} = \vartheta \sigma^i r_j$ ,  $X_i = \vartheta \Phi r_i$ , for the geometrically nonlinear theory of non-shallow shells from (8) and (9) we obtain the following complex form both for the system of equations of equilibrium and Hooke's law:

$$\frac{1}{\Lambda} \frac{\partial \Lambda \tau^{+} \mathbf{r}_{+}}{\partial z} + \frac{\partial \bar{\tau}^{+} \mathbf{r}_{+}}{\partial \bar{z}} - \Lambda (H\tau_{3}^{+} + Q\bar{\tau}_{3}^{+}) + \frac{\partial \tau_{+}^{3}}{\partial x_{3}} + X_{+} = 0,$$

$$\frac{1}{\Lambda} \left( \frac{\partial \Lambda \tau_{3}^{+}}{\partial z} + \frac{\partial \Lambda \bar{\tau}_{3}^{+}}{\partial \bar{z}} \right) + H(\tau_{1}^{1} + \tau_{2}^{2}) + Re\left(\bar{Q}\tau^{+}\mathbf{r}_{+}\right) + \frac{\partial \tau_{3}^{3}}{\partial x_{3}} + X_{3} = 0,$$

$$\tau^{+} = \vartheta \left\{ \lambda \Theta + \mu \left[ \mathbf{R}^{+} \partial_{z} U + \left( \bar{\mathbf{R}}_{+} + 2\partial_{z} U \right) \partial^{\bar{z}} U \right] \right\} \left( \mathbf{R}^{+} + 2\partial^{\bar{z}} U \right)$$

$$+ \mu \vartheta \left\{ \left[ \mathbf{R}^{+} \partial_{\bar{z}} U + \left( \mathbf{R}_{+} + 2\partial_{\bar{z}} U \right) \partial^{\bar{z}} U \right] \left( \bar{\mathbf{R}}^{+} + 2\partial^{\bar{z}} U \right)$$

$$+ \left[ \mathbf{R}^{+} \partial_{3} U + \left( \mathbf{n} + \partial_{3} U \right) \partial^{\bar{z}} U \right] \left( \mathbf{n} + \partial_{3} U \right) \right]$$

$$\tau^{3} = \vartheta \left\{ \left[ \lambda \Theta + 2\mu (\mathbf{n} \partial^{3} U + \frac{1}{2} \partial_{3} U \partial^{3} U \right] \left( \mathbf{n} + \partial_{3} U \right)$$

$$+ \mu \left[ \left( \frac{1}{2} \bar{\mathbf{R}}_{+} \partial_{3} U + \mathbf{n} \partial_{z} U + \partial_{z} U \partial_{3} U \right) \left( \mathbf{R}^{+} + 2\partial^{\bar{z}} U \right)$$

$$+ \left( \frac{1}{2} \mathbf{R}_{+} \partial_{3} U + \mathbf{n} \partial_{\bar{z}} U + \partial_{\bar{z}} U \partial_{3} U \right) \left( \bar{\mathbf{R}}^{+} + 2\partial^{\bar{z}} U \right) \right] \right\},$$

$$(11)$$

$$\left( \partial^{3} U = \partial_{3} U \right).$$

Here

$$\begin{aligned} \boldsymbol{\tau}^{+}\boldsymbol{r}_{+} &= \left(\boldsymbol{\tau}^{1} + i\boldsymbol{\tau}^{2}\right)(\boldsymbol{r}_{1} + i\boldsymbol{r}_{2}), \ \bar{\boldsymbol{\tau}}^{+}\boldsymbol{r}_{+} &= \left(\boldsymbol{\tau}^{1} - i\boldsymbol{\tau}^{2}\right)(\boldsymbol{r}_{1} + i\boldsymbol{r}_{2}), \ \boldsymbol{\tau}_{.3}^{+.} &= \boldsymbol{\tau}^{+}\boldsymbol{n}, \\ \boldsymbol{\tau}_{.+}^{3.} &= \boldsymbol{\tau}^{3}\boldsymbol{r}_{+}, \ \boldsymbol{\tau}_{3}^{3} &= \boldsymbol{\tau}^{3}\boldsymbol{n}, \ 2\partial^{\bar{z}}\boldsymbol{U} &= \left(\boldsymbol{R}^{+}\boldsymbol{R}^{+}\right)\partial_{z}\boldsymbol{U} + \left(\boldsymbol{R}^{+}\bar{\boldsymbol{R}}^{+}\right)\partial_{\bar{z}}\boldsymbol{U}, \ \boldsymbol{\tau}_{.\beta}^{\alpha.} &= \boldsymbol{\tau}^{\alpha}\boldsymbol{r}_{\beta}, \\ \Theta &= 2Re\left[\left(\boldsymbol{R}^{+} + \partial^{\bar{z}}\boldsymbol{U}\right]\right)\partial_{z}\boldsymbol{U} + \partial_{3}\boldsymbol{U}_{3} + 0.5\left(\partial_{3}\boldsymbol{U}\right)^{2}, \ \boldsymbol{R}^{+} &= \vartheta^{-1}\left[\left(1 - Hx_{3}\right)\boldsymbol{r}^{+} + \\ &+ x_{3}Q\bar{\boldsymbol{r}}^{+}\right], \ \boldsymbol{R}_{+} &= (1 - Hx_{3})\boldsymbol{r}_{+} - x_{3}Q\bar{\boldsymbol{r}}_{+}, \ \boldsymbol{R}_{+} &= \boldsymbol{R}_{1} + i\boldsymbol{R}_{2}, \ \boldsymbol{R}^{+} &= \boldsymbol{R}^{1} + i\boldsymbol{R}^{2}, \\ \boldsymbol{R}^{+}\boldsymbol{R}^{+} &= 4x_{3}(\Lambda\vartheta^{2})^{-1}(1 - Hx_{3})Q, \ \boldsymbol{R}^{+}\bar{\boldsymbol{R}}^{+} &= 2(\Lambda\vartheta^{2})^{-1}(\vartheta + 2x_{3}^{2}Q\bar{Q}), \\ \boldsymbol{R}^{+}\boldsymbol{r}_{+} &= 2\vartheta^{-1}Qx_{3}, \ \bar{\boldsymbol{R}}^{+}\boldsymbol{r}_{+} &= 2\vartheta^{-1}(1 - Hx_{3}), \ \boldsymbol{r}^{+}\bar{\boldsymbol{r}}^{+} &= 2\Lambda^{-1}, \ \boldsymbol{r}^{+} &= \boldsymbol{r}^{1} + i\boldsymbol{r}^{2}. \end{aligned}$$

We have the formulas

$$\begin{aligned} \boldsymbol{r}^+ \partial_z \boldsymbol{U} &= \Lambda^{-1} \partial_z U_+ - H U_3, \ \boldsymbol{r}^+ \partial_{\bar{z}} \boldsymbol{U} &= \partial_{\bar{z}} U^+ - Q U_3, \ X_+ = X_1 + i X_2, \\ \boldsymbol{n} \partial_z \boldsymbol{U} &= \partial_z U_3 + 0.5 \left( \bar{Q} U_+ + H \bar{U}_+ \right) \ \left( U^+ = \boldsymbol{U} \boldsymbol{r}^+, \ U_+ = \boldsymbol{U} \boldsymbol{r}_+, \ U_3 = \boldsymbol{U} \boldsymbol{n} \right). \end{aligned}$$

4. In the present paper the three-dimensional problems of the theory of elasticity are reduced to the two-dimensional ones by the method suggested by I. Vekua. Since the system of Legendre polynomials  $\left\{P_m\left(\frac{x_3}{h}\right)\right\}$  is complete

in the interval [-h,h], for equation (8) we obtain the infinite system of twodimensional equations

$$\int_{-h}^{h} \left[ \frac{1}{\sqrt{a}} \frac{\partial \sqrt{a} \vartheta \boldsymbol{\sigma}^{\alpha}}{\partial x^{\alpha}} + \frac{\partial \vartheta \boldsymbol{\sigma}^{3}}{\partial x^{3}} + \vartheta \Phi \right] P_{m} \begin{pmatrix} x_{3} \\ \bar{h} \end{pmatrix} dx_{3} = 0 \quad (m = 0, 1, ...)$$

or in a form

$$\nabla_{\alpha} \overset{(m)}{\sigma}{}^{\alpha} - \frac{2m+1}{h} \left( \overset{(m-1)}{\sigma}{}^{3} + \overset{(m-3)}{\sigma}{}^{3} + \dots \right) + \overset{(m)}{F} = 0, \quad (12)$$

where

$$\begin{pmatrix} {}^{(m)}_{\sigma}{}^{i}, {}^{(m)}_{\Phi} \end{pmatrix} = \frac{2m+1}{2h} \int_{-h}^{h} (\vartheta \sigma^{i}, \vartheta \Phi) P_{m} \begin{pmatrix} {}^{x_{3}}_{\bar{h}} \end{pmatrix} dx_{3} = \frac{2m+1}{2h} \int_{-h}^{h} (\tau^{i}, \mathbf{X}) P_{m} dx_{3},$$

$$\begin{pmatrix} {}^{(m)}_{\mathbf{F}} = {}^{(m)}_{\Phi} + \frac{2m+1}{2h} \begin{pmatrix} {}^{(+)}_{\vartheta} {}^{(+)}_{\sigma} - (-1)^{m} {}^{(-)}_{\vartheta} {}^{(-)}_{\sigma} \end{pmatrix} \begin{pmatrix} {}^{(\pm)}_{\vartheta} = \vartheta(\pm h) \end{pmatrix},$$

 $\nabla_{\alpha}$  are covariant derivatives on the midsurface S.

The equation of state (9) may be written as

where

$$D_{i} \overset{(m)}{U} = \delta_{i}^{\beta} \partial_{\beta} \overset{(m)}{U} + \delta_{i}^{3} \overset{(m)}{U}'; \quad \overset{(m)}{U}' = \frac{2m+1}{h} \begin{pmatrix} {}^{(m+1)} + {}^{(m+3)} + ... \end{pmatrix}, \quad (14)$$

$$\overset{(m)}{\underset{(m_1,m_2,m_3)}{\overset{ijpq}{n_1j_1p_1q_1}}} = \frac{2m+1}{2h} \int_{-h}^{h} \vartheta A_{i_1}^i, A_{j_1}^j, A_{p_1}^p A_{q_1}^q P_{m_1} P_{m_2} P_{m_3} P_m dx_3.$$

The boundary conditions on the lateral contour  $\partial S$  take the form:

a) for the stresses

$$\overset{(m)}{\boldsymbol{\sigma}}_{(l)} = \overset{(m)}{\sigma}_{(ll)}\boldsymbol{l} + \overset{(m)}{\sigma}_{(ls)}s + \overset{(m)}{\sigma}_{(ln)}\boldsymbol{n} = \frac{2m+1}{2h}\int_{h}^{h}\boldsymbol{\sigma}_{(l)}\frac{d\hat{s}}{ds}P_{m}\left(\frac{x_{3}}{h}\right)dx_{3}, \quad (16)$$

b) for the displacements

$$\overset{m}{U} = \frac{2m+1}{2h} \int_{h}^{h} U P_m\left(\frac{x_3}{h}\right) dx_3 = \overset{(m)}{U}_{(l)} l + \overset{(m)}{U}_{(s)}(s) + \overset{(m)}{U}_3 n.$$
(17)

Thus we have constructed an infinite system of two-dimensional equations of geometrically non-linear and non-shallow shells (12-17), which is consistent with the boundary conditions on the face surfaces, i.e.  $\overset{(\pm)}{\sigma}{}^3 = \sigma^3(x^1, x^2, \pm h)$ .

The passage to finite systems can be realized by various methods one of which consists in considering of a finite series, i.e.

$$(\vartheta \boldsymbol{\sigma}^{i}, \boldsymbol{U}, \vartheta \boldsymbol{\Phi}) = \sum_{m=0}^{N} \begin{pmatrix} {}^{(m)}{}_{i}, {}^{(m)}{}_{i}, {}^{(m)}{}_{i} \end{pmatrix} P_{m}\left(\frac{x_{3}}{h}\right) = (\boldsymbol{\tau}^{i}, \boldsymbol{U}, \boldsymbol{X})$$

where N is a fixed nonnegative number. In other words, it is assumed that

$$\overset{(m)}{\boldsymbol{U}} = 0, \quad \overset{(m)}{\boldsymbol{\sigma}}{}^{i} = 0, \quad if \quad m > N.$$

Approximation of this type will be called approximation of order N. The integrals of type (15) can be calculated [5], for example,

$$\binom{m}{A}_{(m_{1})}^{\alpha\beta}{}_{\alpha_{1}\beta_{1}}^{\alpha} = \frac{2m+1}{2h} \int_{-h}^{h} \vartheta^{-1} B^{\alpha}_{\alpha_{1}}(x_{3}) B^{\beta}_{\beta_{1}}(x_{3}) P_{m_{1}}\left(\frac{x_{3}}{h}\right) P_{m}\left(\frac{x_{3}}{h}\right) dx_{3} = \frac{2m+1}{2\sqrt{Eh}} \left[ B^{\alpha}_{\alpha_{1}}(hy) B^{\beta}_{\beta_{1}}(hy) \left( \begin{array}{c} P_{m_{1}}(y) Q_{m}(y), \ m_{1} \leq m \\ Q_{m_{1}}(y) P_{m}(y), \ m_{1} \leq m \end{array} \right) \right]_{y_{1}}^{y_{2}} + \frac{L^{\alpha}_{\alpha_{1}} L^{\beta}_{\beta_{1}}}{K} \sigma^{m}_{m_{1}}, \quad (18)$$

if  $E \neq 0$   $K \neq 0$  and  $a^{\alpha}_{\alpha_1} a^{\beta}_{\beta_1} \delta^m_{m_1}$ , if  $E = H^2 - K = 0$ ; where  $Q_m(y)$  is the Legendre function of the second kind, E is the Euler difference,  $B^{\alpha}_{\beta}(x) = a^{\alpha}_{\beta} + xL^{\alpha}_{\beta}$ ,  $L^{\alpha}_{\beta} = b^{\alpha}_{\beta} - 2Ha^{\alpha}_{\beta}$ . Under the square brackets we mean the following:

$$[f(y)]_{y_1}^{y_2} = f(y_2) - f(y_1), \quad y_{1,2} = [(H \mp \sqrt{E})h]^{-1}.$$

For the integrals containing the product of three (four) Legendre polynomials we have

$$\overset{(m)}{\underset{(m_1,m_2)}{A}} \overset{\alpha_1\alpha_2\alpha_3}{}_{\beta_1\beta_2\beta_3} = \frac{2m+1}{2n} \int\limits_{-h}^{h} \frac{B^{\alpha_1}_{\beta_1}B^{\alpha}_{\beta_2}B^{\alpha}_{\beta_3}}{1-2Hx_3+Kx_3} P_{m_1}P_{m_2}P_m dx_3 = \frac{2m+1}{K^2h^4} \times$$

$$\times \sum_{r=0}^{\min(m_1,m_2)} \alpha_{m_1m_2r} \sum_{n=0}^{3} \mathbb{C}^{\alpha_1\alpha_2\alpha_3}_{\beta_1\beta_2\beta_3} h^n \frac{\partial^2}{\partial y_1 \partial y_2} \left[ \frac{y^n}{y_1 - y_2} \left( \begin{array}{c} P_s(y)Q_m(y), s \le m \\ Q_s(y)P_m(y), s \ge m \end{array} \right) \right]_{y_1}^{y_2}$$

where  $s = m_1 + m_2 - 2r$ ,

$$a_{pqr} = \frac{A_{p-r}A_rA_{q-r}}{A_{p+q-r}} \frac{2(p+q) - 4r + 1}{2(p+q) - 2r + 1}, \quad A_p = \frac{1.3 \cdots 2p - 1}{p!},$$
$$B_{\beta_1}^{\alpha_1}(x)B_{\beta_2}^{\alpha_2}(x)B_{\beta_3}^{\alpha_3}(x) = \sum_{n=0}^3 \mathop{\mathbb{C}_{\beta_1\beta_2\beta_3}^{\alpha_1\alpha_2\alpha_3}} x^n,$$

5. Three-dimensional shell-type bodies are characterized by inequalities

$$|hb^{\alpha}_{\beta}| \leq q < 1 \quad (\alpha,\beta=1,2)$$

Therefore they can be represented as follows

$$|\varepsilon b^{\alpha}_{\beta}R| \le q < 1,$$

where  $\varepsilon = hR^{-1}$  is a small parameter.

Here h is semi-thickness of the shell, R is a certain characteristic radius of curvature of the midsurface S [4].

Now, following Signorini [3] we assume the validity of the expansions

$$\begin{pmatrix} {}^{(m)}{\boldsymbol{\sigma}}{}^{i}, \stackrel{(m)}{\mathbf{U}}, \stackrel{(m)}{\mathbf{F}} \end{pmatrix} = \sum_{n=1}^{\infty} \begin{pmatrix} {}^{(m,n)}{\boldsymbol{\sigma}}{}^{i}, \stackrel{(m,n)}{\mathbf{U}}, \stackrel{(m,n)}{\mathbf{F}} \end{pmatrix} \varepsilon^{n}.$$

Substituting the above expansions into the (12,13) and (10,11) than equalizing the coefficients of expansions for  $\varepsilon^n$  we obtain the following 2-D finite system of equilibrium equations with respect to components of displacement vector in the isometric coordinates  $a_{11} = a_{22} = \Lambda(x^1, x^2)$ , which has the form:

$$4\mu \partial_{\overline{z}} \left( \Lambda^{-1} \partial_{z}^{(m,n)} \right) + 2(\lambda + \mu) \partial_{\overline{z}}^{(m,n)} + 2\lambda \partial_{\overline{z}}^{(m,n)} - (2m+1)\mu \left[ 2\partial_{\overline{z}} \left( \begin{matrix} m-1,n \\ u_{3} \end{matrix} + \begin{matrix} m-3,n \\ u_{3} \end{matrix} + \dotsb \end{matrix} \right) + \begin{matrix} m-1,n \\ u_{+} \end{matrix} + \begin{matrix} m-3,n \\ u_{+} \end{matrix} + \dotsb \end{matrix} + \dotsb \right] + \begin{matrix} m,n \\ F_{+} \end{matrix} = 0, (19) \mu \left( \nabla^{2} \begin{matrix} m,n \\ u_{3} \end{matrix} + \begin{matrix} m,n \\ \theta' \end{matrix} \right) - (2m+1) \left[ \lambda \left( \begin{matrix} m-1,n \\ \theta \end{matrix} + \begin{matrix} m-3,n \\ \theta \end{matrix} + \dotsb \right) + \end{matrix} \\ (\lambda + 2\mu) \left( \begin{matrix} m-1,n \\ u_{3} \end{matrix} + \begin{matrix} m-3,n \\ u_{3} \end{matrix} + \dotsb \right) \right] + \begin{matrix} m,n \\ F_{3} \end{matrix} = 0, \end{cases}$$

where  $u_{+} = u_{1} + iu_{2}$ ,  $\theta = \Lambda^{-1} \Big( \partial_{z} u_{+} \partial_{\overline{z}} \overline{u}_{+} \Big)$ ,  $z = x^{1} + ix^{2}, 2\partial_{z} = \partial_{1} - i\partial_{2}$ ,  $\nabla^{2} = \frac{4}{\Lambda} \frac{\partial^{2}}{\partial z \partial \overline{z}}$ . Obviously, in passing from the *n*-th step of approximation to the (n+1)-th step only the right-hand side of equations are changed. Below we will omit upper index n.

Consider now the cases: N = 0, 1, 2, 3 [6].

Case N = 0. From (3) we get

$$4\mu \partial_{\overline{z}} \left( \Lambda^{-1} \partial_{z} \overset{(0)}{u_{+}} \right) + 2(\lambda + \mu) \partial_{\overline{z}} \overset{(0)}{\theta} = 0,$$
  
$$\mu \nabla^{2} \overset{(0)}{u_{3}} = 0 \qquad \left( \overset{(m)}{F_{+}} = \overset{(m)}{F_{3}} = 0, \ m = 0, 1, 2, 3 \right).$$
 (20)

The complex representation of general solutions has the form

where f(z),  $\varphi(z)$  and  $\psi(z)$  are holomorphic functions of z. We note that for plane (i.e.  $\Lambda = 1$ ) the expression of  $u_{+}^{(0)}$  coincides with the well-known representation of Kolosov-Muskhelishvili.

**Case** N = 1. With respect to the components  $(u_{+}^{(0)}, u_{3}^{(1)})$  and  $(u_{1}^{(1)}, u_{3}^{(0)})$  we have two systems of equations:

$$4\mu\partial_{\overline{z}}\left(\Lambda^{-1}\partial_{z}\overset{(0)}{u_{+}}\right) + 2(\lambda+\mu)\partial_{\overline{z}}\overset{(0)}{\theta} + 2\lambda\partial_{\overline{z}}\overset{(1)}{u_{3}} = 0,$$
  
$$\mu\nabla^{2}\overset{(1)}{u_{3}} + 3\left[\lambda\overset{(0)}{\theta} + (\lambda+2\mu)\overset{(1)}{\theta}\right] = 0$$
(22)

and

$$4\mu\partial_{\overline{z}}\left(\Lambda^{-1}\partial_{z} \overset{(1)}{u_{+}}\right) + 2(\lambda+\mu)\partial_{\overline{z}} \overset{(1)}{\theta} - 3\mu\left(2\partial_{\overline{z}} \overset{(0)}{u_{3}} + \overset{(1)}{u_{+}}\right) = 0,$$

$$\nabla^{2} \overset{(0)}{u_{3}} + \overset{(1)}{\theta} = 0.$$
(23)

The complex representation of general solutions has the form:

and

where  $\Phi(z)$  and  $\Psi(z)$  are holomorphic functions of z.

Note that the systems (10) and (11) coincide with I. Vekua's refined systems of equations for the stretch-strain and bending of plate, respectively.

**Case** N = 2. In this case with respect to the components  $\begin{pmatrix} 0 & 2 & 1 \\ u_1, u_+, u_3 \end{pmatrix}$  and  $\begin{pmatrix} 1 & 0 & 2 \\ u_+, u_3, u_3 \end{pmatrix}$  we have two systems of equations:

$$4\mu \partial_{\overline{z}} \left( \Lambda^{-1} \partial_{z} \overset{(0)}{u_{+}} \right) + 2(\lambda + \mu) \partial_{\overline{z}} \overset{(0)}{\theta} + 2\lambda \partial_{\overline{z}} \overset{(1)}{u_{3}} = 0,$$
  

$$4\mu \partial_{\overline{z}} \left( \Lambda^{-1} \partial_{z} \overset{(2)}{u_{+}} \right) + 2(\lambda + \mu) \partial_{\overline{z}} \overset{(2)}{\theta} - 5\mu \left( 2\partial_{\overline{z}} \overset{(1)}{u_{3}} + 3\overset{(2)}{u_{+}} \right) = 0,$$
  

$$\mu \left( \nabla^{2} \overset{(1)}{u_{3}} + 3\overset{(2)}{\theta} \right) - 3 \left[ \lambda \overset{(0)}{\theta} + (\lambda + 2\mu) \overset{(1)}{u_{3}} \right] = 0$$
(25)

and

$$4\mu \partial_{\overline{z}} \left( \Lambda^{-1} \partial_{z} \overset{(1)}{u} \right) + 2(\lambda + \mu) \partial_{\overline{z}} \overset{(1)}{\theta} - 3\mu \left( 2 \partial_{\overline{z}} \overset{(0)}{u_{3}} + \overset{(1)}{u_{+}} \right) = 0,$$

$$\nabla^{2} \overset{(0)}{u_{3}} + \overset{(1)}{\theta} = 0,$$

$$\mu \nabla^{2} \overset{(2)}{u_{3}} - 5 \left[ \lambda \overset{(1)}{\theta} + 3(\lambda + 2\mu) \overset{(2)}{u_{3}} \right] = 0.$$
(26)

In this case the complex representations of the general solutions of the equations (12) and (13) take the forms

where

$$\nabla^2 v_k = \alpha_k v_k, \ \ \alpha_k^2 - \frac{12(\lambda + \mu)}{\lambda + 2\mu} \alpha_k + \frac{180\mu(\lambda + \mu)}{(\lambda + 2\mu)^2} = 0 \ \ (k = 1, 2); \ \ \nabla^2 \omega = 15\omega,$$

and

$$\begin{aligned}
\overset{(1)}{u_{+}} &= -\frac{1}{\pi} \iint_{S} \frac{\Phi'(\zeta) + \Phi'(\zeta)}{\overline{\zeta} - \overline{z}} dS + \frac{16(\lambda + \mu)}{3(\lambda + 2\mu)} \overline{\Phi''(z)} - \\
&- 2\overline{\Psi'(z)} + i\frac{\partial\chi}{\partial\overline{z}} - \frac{\lambda}{10(\lambda + \mu)} \frac{\partial w}{\partial\overline{z}}, \\
\overset{(0)}{u_{3}} &= \Psi(z) + \overline{\Psi(z)} - \frac{1}{\pi} \iint_{S} \left( \Phi'(\zeta) + \overline{\Phi'(\zeta)} \right) \ln |\zeta - z| dS + \frac{\lambda}{20(\lambda + \mu)} w, \\
\overset{(2)}{u_{3}} &= w - \frac{2\lambda}{3\lambda + 2\mu} \left( \Phi'(z) + \overline{\Phi'(z)} \right),
\end{aligned}$$
(28)

where  $\nabla^2 w = \frac{60(\lambda+\mu)}{\lambda+2\mu} w$ ,  $\nabla^2 \chi = 3\chi$ . **Case** N = 3. For this case we have

$$\begin{aligned}
\overset{(2)}{u_{+}} &= \frac{2}{3} \left( i \frac{\partial \omega}{\partial \overline{z}} + \frac{2\lambda}{3\lambda + 2\mu} \overline{\varphi''} + \sum_{k=1}^{3} \overset{(2)}{A_{k}} \frac{\partial v_{k}}{\partial \overline{z}} \right), \quad (29) \\
\overset{(1)}{u_{3}} &= \sum_{k=1}^{3} \overset{(1)}{A_{k}} v_{k} - \frac{2\lambda}{3\lambda + 2\mu} (\varphi' + \overline{\varphi'}), \\
\overset{(3)}{u_{3}} &= \sum_{k=1}^{3} v_{k}, \quad \left( \nabla^{2} v_{k} = \alpha_{k} v_{k}, \quad \nabla^{2} \omega = 15 \omega \right),
\end{aligned}$$

where

$$\alpha_k^3 - \frac{180(\lambda+\mu)}{\lambda+2\mu} \alpha_k^2 + \frac{120(\lambda+\mu)(7\lambda+15\mu)}{(\lambda+2\mu)^2} \alpha_k + \frac{7 \cdot 900(\lambda+\mu)}{\lambda+2\mu} = 0,$$

and

The general solution of the homogeneous system (19) we can find the form

where  $\varphi'_0(z), \varphi'_1(z), \psi'_0(z), \psi'_1(z)$  are holomorphic functions of z and express the biharmonic solution of the system (19). Then  $\mathfrak{X}_1, \mathfrak{X}_2, \eta_1, \eta_2$  are known constants.

Substituting expressions (31) into (19) the matrix equations for  $\stackrel{(m)}{V_i}$  are obtained

$$\nabla^2 V - AV = X, \quad \nabla^2 \Omega - B\Omega = Y, \tag{32}$$

where V and  $\Omega$  are column-matrices of the form

$$V = \begin{pmatrix} 0 & (1) & (N) & (0) & (1) & (N) \\ V_1, V_1, V_1, V_3, V_3, \dots, V_3 \end{pmatrix}^T, \qquad \Omega = \begin{pmatrix} 0 & (1) & (N) \\ V_2, V_2, \dots, V_2 \end{pmatrix}^T,$$

and A and B are block-matrices  $2N+2 \times 2N+2$  and  $N+1 \times N+1$  respectively.

Using now the formulae Vekua-Bitsadze for the homogenous matrix equations (32) we obtain the following complex representation of the general solutions

$$\begin{split} V &= 2Re\{\varphi(z) + \frac{A}{4}\int_{z_0}^z\int_{\overline{z}_0}^{\overline{z}}\Lambda(t,\overline{t})R(z,\overline{z},t,\overline{t})\varphi(t)dtd\overline{t}\},\\ \Omega &= 2Re\{f(z) + \frac{B}{4}\int_{z_0}^z\int_{\overline{z}_0}^{\overline{z}}\Lambda(t,\overline{t})r(z,\overline{z},t,\overline{t})f(t)dtd\overline{t}\}, \end{split}$$

or

$$V(z,\overline{z}) = 2Re \Big[ \alpha R(z,\overline{z},z_0,\overline{z}_0)\varphi(z) + \int_{z_0}^{z} \Phi(t)R(z,\overline{z},t,\overline{z}_0)dt \Big],$$

$$\begin{split} \Omega(z,\overline{z}) &= 2Re\Big[\beta r(z,\overline{z},z_0,\overline{z}_0) + \int_{z_0}^{z} \Psi(t)r(z,\overline{z},t,\overline{z}_0)f(t)dt\Big],\\ \Big(\varphi(z) &= \frac{\alpha}{2}\int_{z_0}^{z} \Phi(t)dt, \quad V(z_0,\overline{z}_0) = \alpha, \quad \Phi(t) = \frac{\partial V(z_0,\overline{z}_0)}{\partial z}\Big),\\ \Big(f(z) &= \frac{\beta}{2}\int_{z_0}^{z} \Psi(t)dt, \quad \Omega(z_0,\overline{z}_0) = \beta, \quad \Psi(t) = \frac{\partial \Omega(z_0,\overline{z}_0)}{\partial z}\Big), \end{split}$$

where R and r are the Riemann's matrix functions of the equations (32),  $\varphi(z)$  and f(z) are holomorphic column-matrices:

$$\varphi(z) = (\varphi_0(z), \cdots, \varphi_N(z), \varphi_{N+1}(z), \cdots \varphi_{2N}(z))^T, \quad f(z) = (f_0(z), \cdots, f_N(z))^T.$$

Then particular solutions of the matrix equations (32) have the form

$$\begin{split} & \stackrel{\wedge}{V}(z,\overline{z}) = \frac{1}{4} \int_{z_0}^{z} \int_{\overline{z_0}}^{\overline{z}} \Lambda(t,\overline{t}) R(z,\overline{z},t,\overline{t}) X(t,\overline{t}) dt d\overline{t}, \\ & \stackrel{\wedge}{\Omega}(z,\overline{z}) = \frac{1}{4} \int_{z_0}^{z} \int_{\overline{z_0}}^{\overline{z}} \Lambda(t,\overline{t}) r(z,\overline{z},t,\overline{t}) Y(t,\overline{t}) dt d\overline{t}, \end{split}$$

where

$$\begin{aligned} R(z,\overline{z},t,\overline{t}) &= E + \frac{A}{4} \int_{t}^{z} \int_{\overline{t}}^{\overline{z}} \Lambda(t_{1},\overline{t}_{1}) dt_{1} d\overline{t}_{1} + \\ \left(\frac{A}{4}\right)^{2} \int_{t}^{z} \int_{\overline{t}}^{\overline{z}} \Lambda(t_{1},\overline{t}_{1}) \left(\int_{t}^{t} \int_{\overline{t}}^{\overline{t}} \Lambda(t_{2},\overline{t}_{2}) dt_{2} d\overline{t}_{2}\right) dt_{1} d\overline{t}_{1} \cdots \\ r(z,\overline{z},t,\overline{t}) &= E + \frac{B}{4} \int_{t}^{z} \int_{\overline{t}}^{\overline{z}} \Lambda(t_{1},\overline{t}_{1}) dt_{1} d\overline{t}_{1} + \\ \left(\frac{B}{4}\right)^{2} \int_{t}^{z} \int_{\overline{t}}^{\overline{z}} \Lambda(t_{1},\overline{t}_{1}) \left(\int_{t}^{t} \int_{\overline{t}}^{\overline{t}} \Lambda(t_{2},\overline{t}_{2}) dt_{2} d\overline{t}_{2}\right) dt_{1} d\overline{t}_{1} + \cdots \end{aligned}$$

For the first boundary condition (in stress) we have

$$(\lambda+\mu)^{(m)} - 2\mu\Lambda \frac{\partial^{(m)}_{u_+}}{\partial \overline{z}} \left(\frac{d\overline{z}}{ds}\right)^2 = {}^{(m)}_{a_1} + i{}^{(m)}_{b_1}, \quad Im\left(\frac{\partial^{(m)}_{u_3}}{\partial \overline{z}}\frac{d\overline{z}}{ds}\right) = {}^{(m)}_{c_1} \quad (on \ \partial S)$$

The second boundary condition (in displacements) for any m takes the form

$${}^{(m)}_{u_{+}}\frac{d\overline{z}}{ds} = {}^{(m)}_{a_{2}} + i {}^{(m)}_{b_{2}}, \quad {}^{(m)}_{u_{3}} = {}^{(m)}_{c_{2}} \quad (on \ \partial S).$$

The basic boundary conditions (N = 0) for any *n* have the form: a) for the first boundary problem (in displacements)

$${}^{(0,n)}_{U_{(\ell)}} + i {}^{(0,n)}_{U_{(s)}} = i {}^{(0,n)}_{U_+} \frac{d\bar{z}}{ds} = {}^{(n)}_{d_+}, {}^{(0,n)}_{U_3} = {}^{(n)}_{d_3} \text{ on } \partial D$$
(33)

b) for the second boundary problem (in stresses)

$$\begin{aligned} {}^{(0,n)}_{\sigma_{(\ell\ell)}} + i {}^{(0,n)}_{\sigma_{(\ells)}} &= \frac{1}{2} \begin{bmatrix} {}^{(0,n)}_{\sigma} + \boldsymbol{r}_{+} - \begin{pmatrix} {}^{(0,n)}_{\sigma} + \boldsymbol{r}_{+} \end{pmatrix} \frac{dz}{d\bar{z}} \end{bmatrix} = {}^{(n)}_{e}_{+}, \\ {}^{(0,n)}_{\sigma_{(\elln)}} &= -Im \left[ \begin{pmatrix} {}^{(0,n)}_{\sigma} + \boldsymbol{n} \end{pmatrix} \frac{d\bar{z}}{ds} \right] = {}^{(n)}_{e_{3}} \text{ on } \partial D. \end{aligned}$$

$$(34)$$

Here we present a general scheme of solution of boundary problems when the domain D is a circle of radius  $r_0$ .

The first boundary problem for any n takes the form (on  $|z| = r_0$ ),

where  $\overset{(n)}{G_+}$  and  $\overset{(n)}{G_3}$  are the known values containing solutions  $\overset{(0,1)}{U_i}, \cdots, \overset{(0,n-1)}{U_i}$  of the previous approximations.

Let  $\Lambda(z, \bar{z})$  depend only on r = |z|, next  $\varphi'(z)$ ,  $\psi(z)$  and  $\overset{(n)}{G}_+$  are expanded in power series of the type

$$\varphi'(z) = \sum_{k=0}^{\infty} a_k z^k, \ \Psi(z) = \sum_{k=0}^{\infty} b_k z^k, \ G_+ = \sum_{k=-\infty}^{\infty} A_k e^{ik\vartheta}.$$

Substituting these expansions into (24), we obtain

$$a_{0} = \frac{r_{0}}{\alpha_{0}} \frac{\varpi A_{1} + \bar{A}_{1}}{\varpi^{2} - 1}, \ a_{k} = \frac{r_{0}^{k+1} A_{k+1}}{\varpi \alpha_{k}} \ (k \ge 1),$$
$$b_{k} = -\frac{\bar{A}_{k}}{r_{0}^{k}} - \frac{\alpha_{0} r_{0}^{k+2}}{\varpi \alpha_{k+1}} A_{k+2}, (k \ge 0), \ \alpha_{k} = 2 \int_{0}^{r_{0}} \rho^{2k+1} \Lambda(\rho) d\rho.$$

 $\overset{(0,n)}{U}_{_3}$  is representable in the form of the Poisson integral,

$${}^{(0,n)}_{U_3}(r,\vartheta) = \frac{1}{2\pi} \int_0^{2\pi} {}^{(n)}_{G_3}(\psi) \frac{r_0^2 - r^2}{r^2 - 2r_0 r \cos(\psi - \vartheta) + r_0^2} d\psi.$$
(37)

Thus for any n we can construct formal solutions of the problem (22), when N = 0.

From the second boundary condition (23), we obtain (on  $\partial D$ )

$${}^{(0,n)}_{\sigma_{(\ell\ell)}} + i {}^{(0,n)}_{\sigma_{(\ell s)}} = {}^{(n)}_{\ell_{+}} \Rightarrow (\lambda + \mu) {}^{(0,n)}_{\Theta} - 2\mu \left(\frac{1}{\Lambda} {}^{(0,n)}_{U_{+}}\right) \frac{d\bar{z}}{dz} = {}^{(n)}_{P_{+}}, \qquad (38)$$

$$\overset{(0,n)}{\sigma}_{(\ell n)} = \overset{(n)}{\ell_3} \Rightarrow \operatorname{Im} \left( \frac{\partial \overset{(0,n)}{U_3}}{\partial z} \frac{d\bar{z}}{ds} \right) = \overset{(n)}{P_3},$$

$$\left( \overset{(0,n)}{\Theta} = \frac{1}{\Lambda} \left( \frac{\partial \overset{(0,n)}{U_+}}{\partial z} + \frac{\partial \overset{(0,n)}{\bar{U}_+}}{\partial \bar{z}} \right) \right).$$

$$(39)$$

Consider the case of a spherical shell, whose midsurface is a spherical segment of radius  $R_0 \sin \vartheta$ , where  $R_0$  is the radius of a sphere. Isometric coordinates on the sphere can be represented in the form

$$z = x^{1} + ix^{2} = re^{i\varphi}, \quad r = \operatorname{tg}\frac{\vartheta}{2}, \quad \Lambda = 4R^{2}(1 + z\bar{z})^{-2} \quad (0 \le \vartheta \le \vartheta_{0})$$

Let the expressions

$$(\varphi'(z), \Psi'(z), f(z)) = \sum_{k=0}^{\infty} (a_k, b_k, c_k) z^k, \ \begin{pmatrix} {}^{(n)}_{g_+}, {}^{(n)}_{g_3} \end{pmatrix} = \sum_{k=-\infty}^{\infty} (A_k, B_k) e^{ik\varphi},$$

be valid, where  $\stackrel{(n)}{g_+}$  and  $\stackrel{(n)}{g_3}$  are known values expressed by  $\stackrel{(0,1)}{U_i}, \cdots, \stackrel{(0,n-1)}{U_i}$  of the previous approximations. Substituting these expansions into (27), (28)and taking into account that principal vector and moment of stresses are zero, we obtain

$$a_{k} = \frac{A_{k}}{2\mu r_{0}^{k}} \frac{1}{1+2\omega(1+r_{0}^{2})\beta_{k}},$$
  

$$b_{k} = \frac{-1}{2\mu r_{0}^{k-1}} \frac{(1+r_{0}^{2})^{-1}}{k+2r_{0}^{2}} \left[ \frac{((1+r_{0}^{2})k+2r_{0}^{2})A_{k+1}}{1+2\omega(1+r_{0}^{2})\beta_{k+1}r_{0}^{2}} + \bar{A}_{k-1} \right] \quad (k \ge 0).$$
  

$$c_{k} = \frac{2}{\mu} \frac{R_{0}}{1+r_{0}^{2}} \frac{B_{k}}{kr_{0}^{k-1}} \quad (k \ge 1), \quad B_{0} = 0, \quad \beta_{k}(z) = \frac{1}{z^{k+2}} \int_{0}^{z} \frac{(z-t)t^{k}dt}{(1+t\bar{z})^{3}}.$$

From here we obtain the well-known Dini's formula

$${}^{(0,n)}_{U_3}(r_0,\varphi) = -\frac{r_0}{\pi} \int_{0}^{2\pi} {}^{(n)}_{P_3}(r_0,\varphi) \ln |\sigma - z| d\vartheta + \text{const} \quad (\sigma = r_0 e^{i\vartheta}).$$

## References

[1] I.N. Vekua. Shell Theory: General Methods of Construction. Pitman Advanced Publishing Program, Boston-London-Melburne, 1985, 287p.

[2] A. Bitsadze. Boundary Value Problems for Second Order Elliptic Equations. Moscow, Nauka, 1966, 203p (in Russian)

[3] P.G. Ciarlet. Mathematical Elasticity. V.I three-dimensional elasticity. North-Holland Publishing, Amsterdam-New-Oxford-York-Tokyo, 1998, 471p.

[4] A.L. Goldenveizer. Theory of Elastic thin Shells. Moscow, Nauka, 1976, 512p. (in Russian)

[5] T.V. Meunargia. A small parameter method for I. Vekua's nonlinear and non-shallow shells. Proceedings of the IUTAM Symposium on Relation of Shell, Plate, Beam, and 3D Models Dedicated to Centenary of Ilia Vekua, 23-27 April, 2007, Tbilisi, Georgia. IUTAM Bookseries, 9 (2008), 155-166

[6] T.V. Meunargia. On the complex representations for the nonlinear and non-shallow shells. Rep. Enlarged Sess. Semin. I. Vekua Appl. Math., 23 (2009), 7-11

Received 30.11.2009; revised 23.12.2009; accepted 29.12.2009